Please check the examination deta	nils below before ente	ring your candidate information		
Candidate surname		Other names Britishs		
Pearson Edexcel International Advanced Level	Centre Number	Candidate Number		
<b>Time</b> 1 hour 45 minutes	Paper reference	WPH15/01		
Physics				
International Advanced Level UNIT 5: Thermodynamics, Radiation, Oscillations and Cosmology				

### **Instructions**

- Use **black** ink or ball-point pen.
- **Fill in the boxes** at the top of this page with your name, centre number and candidate number.
- Answer **all** questions.
- Answer the questions in the spaces provided
  - there may be more space than you need.
- Show all your working in calculations and include units where appropriate.

#### Information

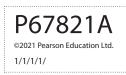
- The total mark for this paper is 90.
- The marks for **each** question are shown in brackets
  - use this as a guide as to how much time to spend on each question.
- In the question marked with an asterisk (\*), marks will be awarded for your ability to structure your answer logically, showing how the points that you make are related or follow on from each other where appropriate.
- The list of data, formulae and relationships is printed at the end of this booklet.

### **Advice**

- Read each question carefully before you start to answer it.
- Try to answer every question.
- Check your answers if you have time at the end.
- Good luck with your examination.

Turn over ▶







SECTION A

Answer ALL the questions in this section.

For questions 1-10, in Section A, select one answer from A to D and put a cross in the box 

Market State of the section A. or questions 1-10, in Section A, select one answer from A to D and put a cross in the work.

If you change your mind, put a line through the box A and then mark your new answer, with a cross A.

			with a cross Z.			
1			tor and counter were placed near a radium source. A student measured the count minute. He used this value to calculate the count rate in Bq from the source.			
	Wh	ich (	of the following would increase the accuracy of the student's value for the count rate?			
	X	A	Adding the background count rate to the measured count rate.			
	X	В	Increasing the counting time to 10 minutes.			
C Repeating the count with a different detector and calculating a mean value.						
	X	D	Subtracting the background count from the measured count.			
			(Total for Question 1 = 1 mark)			
2			ial value determined by astronomers for the Hubble constant was smaller than ently accepted value.			
	Wh	ich (	of the following statements is correct?			
	X	A	The universe is bigger than astronomers initially thought.			
	X	В	The universe is not as old as astronomers initially thought.			
	×	C	The universe is older than astronomers initially thought.			
	X	D	The universe is smaller than astronomers initially thought.			
			(Total for Question 2 = 1 mark)			
3	Wh	en a	train travels across a railway bridge, the bridge may be forced into resonant oscillation.			
	Wh	ich o	of the following happens only when resonance occurs?			
	×	A	Damping reduces the amplitude of oscillation.			
	×	В	Energy is transferred from the train to the bridge.			
	X	C	Resistive forces dissipate energy from the train.			
	×	D	The amplitude of oscillation is unusually large.			

(Total for Question 3 = 1 mark)

4 In the Sun, nuclear fusion occurs mainly in the core.

Which of the following is a reason for this?

- A Helium nuclei surround the core.
- **B** Most of the Sun's hydrogen is in the core.
- C The core contains most of the mass of the Sun.
- **D** The temperature is greatest in the core.

(Total for Question 4 = 1 mark)

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5 The gravitational field strength on the surface of Mercury is  $g_{\rm M}$ . Callisto, a moon of Jupiter, has the same radius as Mercury but only one third of its density.

What is the gravitational field strength on the surface of Callisto?

- $\triangle$  A  $\frac{g_{\rm M}}{q}$
- $\square$  B  $\frac{g_{\rm M}}{3}$
- $\square$  C  $3g_{\rm M}$
- $\square$  **D**  $9g_{\rm M}$

(Total for Question 5 = 1 mark)

6 Tau Ceti is one of the closest stars to the Sun. Tau Ceti has a surface temperature of  $5300\,\mathrm{K}$  and a luminosity of  $0.55\,L_\mathrm{Sun}$ , where  $L_\mathrm{Sun}$  is the luminosity of the Sun.

Which of the following is the correct description of Tau Ceti?

- A main sequence star
- **B** red dwarf star
- C red giant star
- **D** white dwarf star

(Total for Question 6 = 1 mark)

Helium gas in a closed cylinder is heated until the pressure exerted by the helium gas in a closed cylinder is heated until the pressure exerted by the helium gas in a closed cylinder is heated until the pressure exerted by the helium gas in a closed cylinder is heated until the pressure exerted by the helium gas in a closed cylinder is heated until the pressure exerted by the helium gas in a closed cylinder is heated until the pressure exerted by the helium gas in a closed cylinder is heated until the pressure exerted by the helium gas in a closed cylinder is heated until the pressure exerted by the helium gas in a closed cylinder is heated until the pressure exerted by the helium gas in a closed cylinder is heated until the pressure exerted by the helium gas in a closed cylinder is heated until the pressure exerted by the helium gas constants, and the pressure exerted by the helium gas constants,

What is the ratio  $\frac{\langle v_{\rm F}^2 \rangle}{\langle v_{\rm I}^2 \rangle}$ ?

- X **A** 1
- 2
- X 8 D

(Total for Question 7 = 1 mark)

Two stars, P and Q, have approximately the same radius. The surface temperature of star P is twice the surface temperature of star Q.

Star P has a luminosity  $L_p$  and star Q has a luminosity  $L_Q$ .

The ratio  $\frac{L_p}{L_0}$  is approximately

- 2

- **D** 16

(Total for Question 8 = 1 mark)

An object oscillates with simple harmonic motion. The period of oscillation is 0.25 s and the amplitude is 0.25 m.

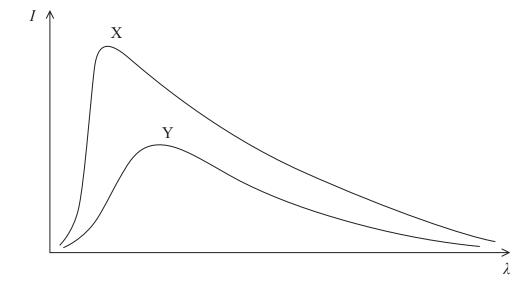
Which of the following shows a correct calculation of the maximum velocity of the object?

North, Roy, Comp. 9

- **B**  $\frac{2\pi}{0.55} \times 0.25$
- C  $(2\pi \times 0.55)^2 \times 0.25$
- $\mathbf{D} \quad \left(\frac{2\pi}{0.55}\right)^2 \times 0.25$

(Total for Question 9 = 1 mark)

10 Two stars, X and Y, behave as black body radiators. The graphs show how the intensity I of radiation emitted varies with wavelength  $\lambda$  for the two stars.



The stars are both viewed from the same distance.

Which of the following statements can be deduced from the two graphs?

- X A Star X is brighter and has a higher surface temperature than star Y.
- X Star X is brighter and has a lower surface temperature than star Y.
- X Star X is dimmer and has a higher surface temperature than star Y.
- X Star X is dimmer and has a lower surface temperature than star Y.

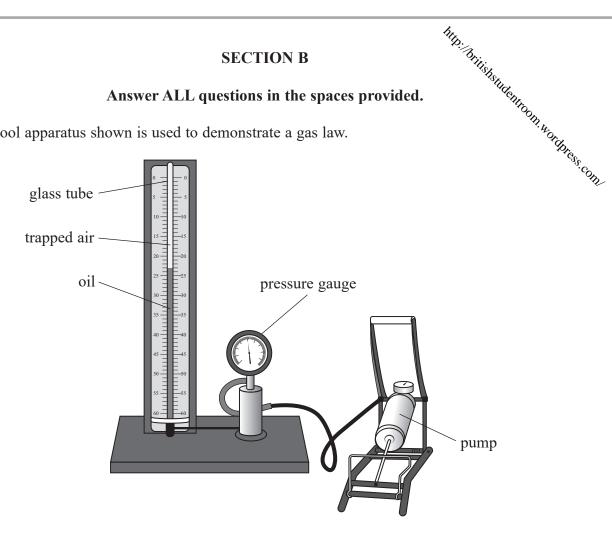
(Total for Question 10 = 1 mark)

**TOTAL FOR SECTION A = 10 MARKS** 

## **SECTION B**

# Answer ALL questions in the spaces provided.

11 The school apparatus shown is used to demonstrate a gas law.



Air is trapped in a glass tube of uniform cross-sectional area. A pump forces oil into the base of the glass tube. This forces the air into a smaller volume. The pressure of the trapped air is displayed on the pressure gauge.

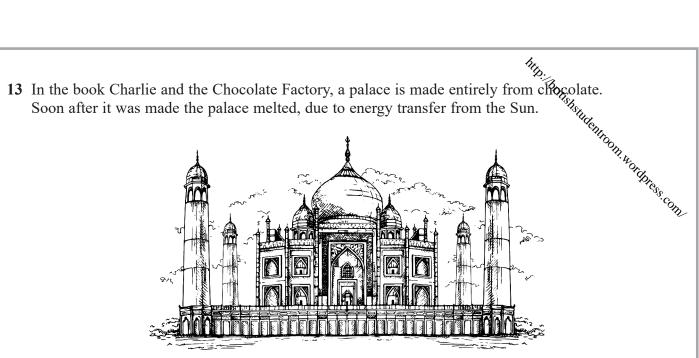
(	(a)	The	pressure	of the	trapped ai	r increases	when	the ai	r is	forced	into a	smaller	volume.

Explain	ı why, usıng	g ideas of i	molecular	motion.

(4)

(b) The apparatus is used in a laboratory where the temperature is 293 K.  When the air occupies a volume of $2.43 \times 10^{-3} \mathrm{m}^3$ the reading on the pressure gaus is $1.05 \times 10^5 \mathrm{Pa}$ .  Calculate the number of molecules of air trapped in the glass tube.	
When the air occupies a volume of $2.43 \times 10^{-3}$ m <sup>3</sup> the reading on the pressure gains $1.05 \times 10^{5}$ Pa.	Ngentroom
Calculate the number of molecules of air trapped in the glass tube.	(2) Cop
Number of molecules of air =	
(Total for Question 11 = 6	marks)

	Very 1 and 1	
(a) S	cent years, astronomers have discovered sources of fast radio bursts (FRBs) in galaxies. Studies suggest that these sources may be a type of standard candle state what is meant by a standard candle.	OOM 40.
		<sup>7</sup> %,
(	An FRB source emits intense bursts of radio waves, each burst lasting for a fraction of a second. The closest FRB source is in a massive spiral galaxy $4.60 \times 10^{24}$ m from the Earth.	
	A detector of area $1.00 \times 10^{-4}  \text{m}^2$ on the surface of the Earth received bursts of radio vaves. In one burst, $9.40 \times 10^{-23}  \text{J}$ of energy was received in a time of 1.15 ms.	
(	i) Show that the luminosity of the source is about $2 \times 10^{35}$ W.	(4)
		(4)
(	ii) When FRB sources were first discovered, some observers suggested that the bursts might be alien communications.	
	Suggest why this is unlikely.	
	luminosity of the Sun = $3.8 \times 10^{26}$ W	(8)
		(2)



(Source: Morkhatenok/Shutterstock)

(Total for Question 13 = 6 marks)

The total volume of chocolate used for the palace was 1250 m<sup>3</sup>. The initial temperature of the chocolate palace was 28.5 °C. Chocolate melts at a temperature of 36.0 °C.

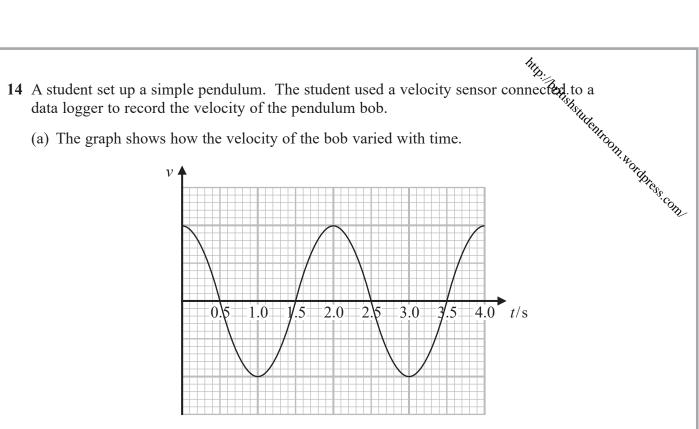
The book states that the palace melted completely in less than a day.

Deduce whether this statement could be correct.

rate of energy transfer to the palace =  $7.5 \times 10^5 \,\mathrm{W}$ density of chocolate =  $1325 \,\mathrm{kg}\,\mathrm{m}^{-3}$ specific latent heat of chocolate =  $4.5 \times 10^4 \, \mathrm{J \, kg^{-1}}$ specific heat capacity of chocolate is  $1.30 \times 10^3 \,\mathrm{Jkg^{-1}\,K^{-1}}$ 



**(6)** 



(i) Add a line to the graph to show how the displacement of the bob varied with time during the same time interval.

**(2)** 

(ii) Determine the length of the simple pendulum.

(3)

Length of simple pendulum =

(b) In some experiments, it is an advantage to use a data logger rather than other measuring instruments.

Describe when the use of a data logger is an advantage.

**(2)** 

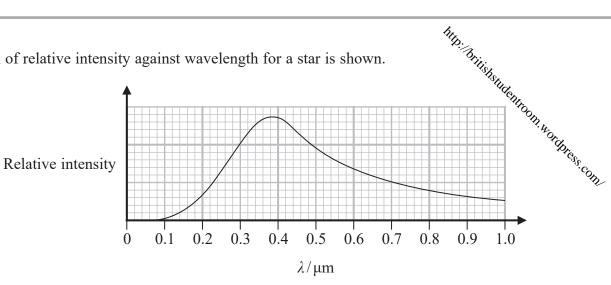
(Total for Question 14 = 7 marks)

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15 A graph of relative intensity against wavelength for a star is shown.

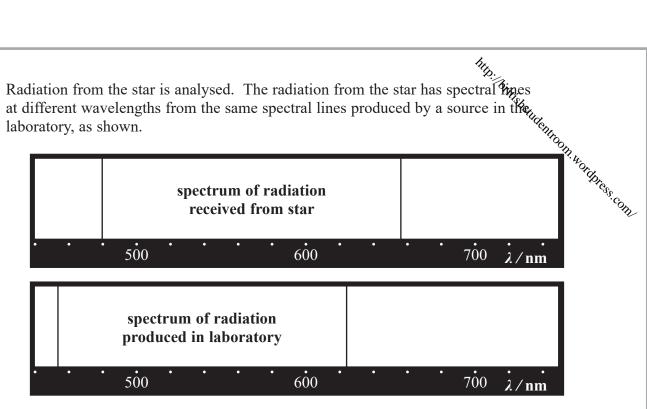


(a) Determine the surface temperature of the star.

(3)

Surface temperature of star =

(b) Radiation from the star is analysed. The radiation from the star has spectral times



Explain what can be concluded about the star from these results. Your answer should include a calculation.

(Total for Questi	ion 15 = 8 marks)

(5)

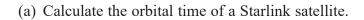
(3)

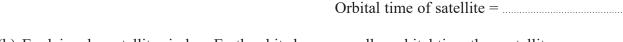
In 2015, the company SpaceX stated its plan to launch about 4,000 "Starlink" satellites into a low Earth orbit. These satellites would provide a low-cost internet service to the people around the world.

200 Starlink satellites had been launched into an orbit

\*\*The company SpaceX stated its plan to launch about 4,000 "Starlink" satellites into a low-cost internet service to the company of the people around the world. 16 In 2015, the company SpaceX stated its plan to launch about 4,000 "Starlink" satellites

radius of Earth =  $6.4 \times 10^6$  m

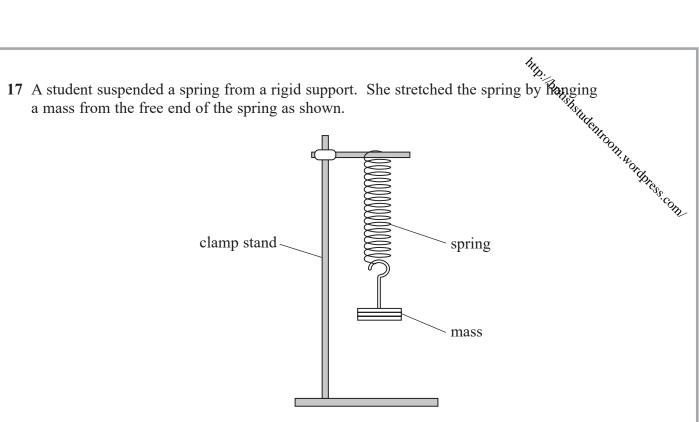




(b) Explain why satellites in low Earth orbits have a smaller orbital time than satellites in higher Earth orbits. (3)



	Aug.
(c) The satellites were launched into orbit using	rockets able to make multiple space flights.
Calculate the minimum kinetic energy require	ed to raise the satellite to its lower orbit height.
mass of a Starlink satellite = 227 kg.	rockets able to make multiple space flights.  ed to raise the satellite to its lower of hit height.  (3) Orthress Com
	Minimum kinetic energy =
	(Total for Question 16 = 9 marks)



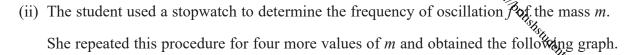
She set the mass into vertical oscillations by displacing it a small distance from its equilibrium position. The spring obeys Hooke's law.

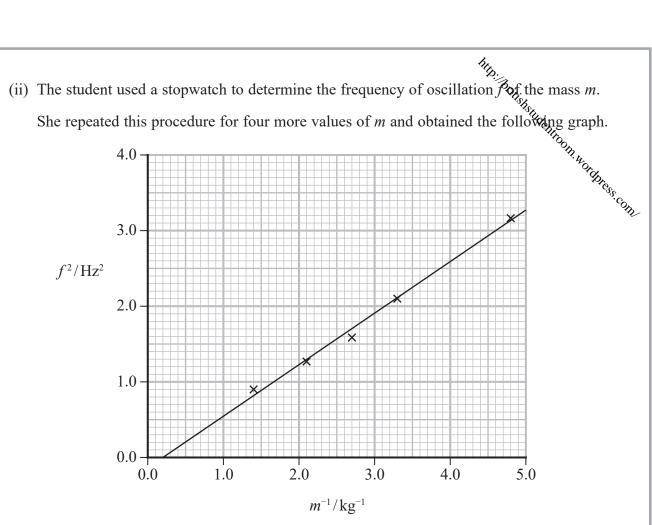
(a) Explain why the mass moves with simple harmonic motion.

C	Z	)	
1		/	

- (b) When the mass on the end of the spring was 0.200 kg, the extension of the spring was 7.50 cm when the mass was in equilibrium.
  - (i) Show that the stiffness k of the spring is about  $26 \,\mathrm{N}\,\mathrm{m}^{-1}$ .







The student used the graph to determine a value for k.

Deduce whether the graph gives a value of k consistent with the value in (i).

(Total for Question 17 = 10 marks)

**(6)** 

Many countries use nuclear power stations to provide electrical power. Energy fixing released when nuclei undergo fission in the core of the reactor.

The provided electrical power is a superior of the reactor.

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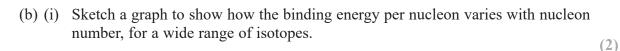
The provided electrical power is a superior of the reactor.

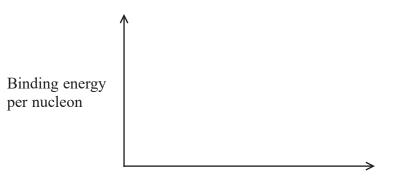
The provided electrical power is a superior of the reactor.

The provided electrical power is a superior of the reactor.

The provided electrical power is a superior of the reactor.

The





Nucleon number

(ii) Mark the position of iron-56 on your graph.

**(1)** 

(c) Complete the nuclear equation below.

$$^{236}_{92}U \rightarrow ^{--}_{38}Sr + ^{-141}Xe + 2 \times ^{--}_{n}$$

**(2)** 

(d) Calculate, in MeV, the binding energy per nucleon for a nucleus of  $^{236}_{92}$ U.

	Mass/GeV/c <sup>2</sup>
<sup>236</sup> U	219.8750
n	0.93956
p	0.93827

**(2)** 

Binding energy per nucleon = ...... MeV

h <sub>th</sub> .
*(e) Ultraviolet radiation (UV) is produced when alpha particles interact with air. This can be used to detect alpha particles when a nuclear reactor is decommissioned.
*(e) Ultraviolet radiation (UV) is produced when alpha particles interact with air.  This can be used to detect alpha particles when a nuclear reactor is decommissioned.  Explain how UV is produced by alpha particles in the air, and why detecting UV has on advantages compared with detecting alpha particles directly.  (6) Thress conn
(6) M <sub>ress</sub> Con
(Total for Question 18 = 14 marks)

**(5)** 

19 A pacemaker is a device used to regulate a person's heart rate.

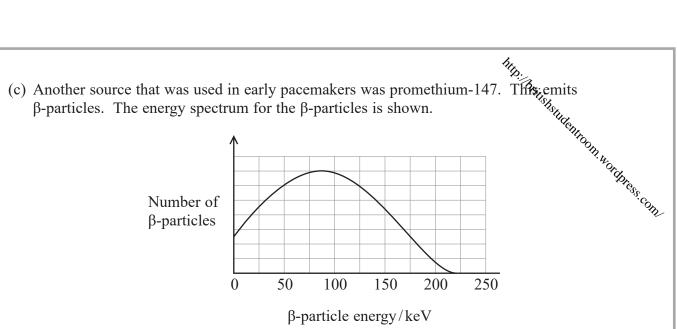
uranium is about 5.6 MeV.

	Mass/u
Plutonium nucleus	237.999089
Uranium nucleus	233.991578
α-particle	4.001506

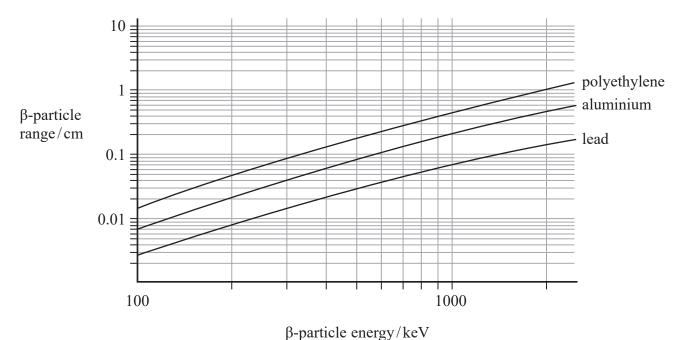

(b)	In one pacemaker, the activity of the plutonium source was measured to be $6.75\times10^{10}\mathrm{Bq}$ in 2020.
	Calculate the power of the source, in W, when it was fitted 40 years ago

(h)	In one pacemaker, the activity of the plutonium source was measured to	
(0)	be $6.75 \times 10^{10}$ Bq in 2020.	
	Calculate the power of the source, in W, when it was fitted 40 years ago.	
	half-life of Pu-238 = 87.7 years	
	In one pacemaker, the activity of the plutonium source was measured to be $6.75 \times 10^{10}$ Bq in 2020.  Calculate the power of the source, in W, when it was fitted 40 years ago. half-life of Pu-238 = 87.7 years  energy of $\alpha$ -particle = 5.6 MeV	-
		•
		•

Power of source =



The graphs below show how the range of beta particles depends on the beta particle energy, for different materials.



It is suggested that a layer of polyethylene, of absorb all the beta particles emitted from a pr	It is suggested that a layer of polyethylene, of thickness 0.5 cm, would be able to absorb all the beta particles emitted from a promethium-147 source.  Comment on this suggestion.			
Comment on this suggestion.	*entroon Andrews			
	**************************************			
	(Total for Question 19 = 13 marks)			

TOTAL FOR SECTION B = 80 MARKS TOTAL FOR PAPER = 90 MARKS



## List of data, formulae and relationships

Acceleration of free fall

close to Earth's surface)

(close to Earth's surface)

(close to Earth's surface)  $g = 9.81 \text{ m s}^{-2}$ 

 $k = 1.38 \times 10^{-23} \text{ J K}^{-1}$ Boltzmann constant

Coulomb's law constant  $k = 1/4\pi\varepsilon_0$ 

 $= 8.99 \times 10^9 \text{ N m}^2 \text{ C}^{-2}$ 

 $e = -1.60 \times 10^{-19}$ C Electron charge

 $m_e = 9.11 \times 10^{-31} \text{kg}$ Electron mass

 $1 \text{ eV} = 1.60 \times 10^{-19} \text{ J}$ Electronvolt

 $G = 6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2}$ Gravitational constant

 $g = 9.81 \text{ N kg}^{-1}$ Gravitational field strength (close to Earth's surface)

 $\varepsilon_0 = 8.85 \times 10^{-12} \text{ F m}^{-1}$ Permittivity of free space

 $h = 6.63 \times 10^{-34} \,\mathrm{J s}$ Planck constant

 $m_p = 1.67 \times 10^{-27} \text{ kg}$ Proton mass

 $c = 3.00 \times 10^8 \,\mathrm{m \ s^{-1}}$ Speed of light in a vacuum

 $\sigma = 5.67 \times 10^{-8} \; W \; m^{-2} \; K^{-4}$ Stefan-Boltzmann constant

 $u = 1.66 \times 10^{-27} \text{ kg}$ Unified atomic mass unit

### Unit 1

Mechanics

Kinematic equations of motion 
$$s = \frac{(u+v)t}{2}$$

$$v = u + at$$

$$s = ut + \frac{1}{2}at^2$$

$$v^2 = u^2 + 2as$$

 $\Sigma F = ma$ Forces

$$g = \frac{F}{m}$$

$$W = mg$$

Momentum p = mv

Moment of force moment = Fx

 $\Delta W = F \Delta s$ Work and energy

$$E_{\rm k} = \frac{1}{2} \, m v^2$$

 $\Delta E_{\rm grav} = mg\Delta h$ 

 $P = \frac{E}{t}$ 

 $P = \frac{W}{t}$ 

Power



Efficiency

$$efficiency = \frac{useful energy output}{total energy input}$$

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Materials

Density

 $\rho = \frac{m}{V}$  $F = 6\pi \eta r v$ Stokes' law

 $\Delta F = k\Delta x$ Hooke's law

 $\Delta E_{\rm el} = \frac{1}{2} F \Delta x$ Elastic strain energy

 $E = \frac{\sigma}{\varepsilon}$  where Young modulus

Stress  $\sigma = \frac{F}{A}$ 

Strain  $\varepsilon = \frac{\Delta x}{x}$ 



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### Unit 2

#### Waves

Wave speed  $v = f\lambda$ Speed of a transverse wave on a string  $v = \sqrt{\frac{T}{\mu}}$ 

Intensity of radiation  $I = \frac{P}{A}$ 

Refractive index  $n_1 \sin \theta_1 = n_2 \sin \theta_2$ 

$$n=\frac{c}{v}$$

Critical angle  $\sin C = \frac{1}{n}$ 

Diffraction grating  $n\lambda = d\sin\theta$ 

# Electricity

Potential difference  $V = \frac{W}{Q}$ 

Resistance  $R = \frac{V}{I}$ 

Electrical power, energy P = VI

 $P = I^2 R$ 

 $P = \frac{V^2}{R}$ 

W = VIt

Resistivity  $R = \frac{\rho l}{A}$ 

Current  $I = \frac{\Delta Q}{\Delta t}$ 

I = nqvA

Resistors in series  $R = R_1 + R_2 + R_3$ 

Resistors in parallel  $\frac{1}{R} = \frac{1}{R_1} + \frac{1}{R_2} + \frac{1}{R_3}$ 

## Particle nature of light

Photon model E = hf

Einstein's photoelectric  $hf = \emptyset + \frac{1}{2} mv_{\text{max}}^2$  equation

de Broglie wavelength  $\lambda = \frac{h}{p}$ 



## Unit 4

Mechanics

Impulse

Kinetic energy of a non-relativistic particle

motion in a circle

$$F\Delta t = \Delta p$$

 $E_k = \frac{p^2}{2m}$ 

 $v = \omega r$ 

$$T = \frac{2\pi}{\omega}$$

$$a = \frac{v^2}{r}$$

Centripetal force

$$F = ma = \frac{mv^2}{r}$$

$$F = m\omega^2 r$$

Electric and magnetic fields

Electric field

$$E = \frac{F}{Q}$$

Coulomb's law

$$F = \frac{Q_1 Q_2}{4\pi \varepsilon_0 r^2}$$

$$E = \frac{Q}{4\pi\varepsilon_0 r^2}$$

$$E = \frac{V}{d}$$

**Electrical Potential** 

$$V = \frac{Q}{4\pi\varepsilon_0 r}$$

Capacitance

$$C = \frac{Q}{V}$$

Energy stored in capacitor

$$W = \frac{1}{2}QV$$

$$W = \frac{1}{2}CV^2$$

$$W = \frac{1}{2} \frac{Q^2}{C}$$

Capacitor discharge

$$Q = Q_0 e^{-t/RC}$$



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Resistor capacitor discharge

$$I = I_0 \mathrm{e}^{-t/RC}$$

$$V = V_0 e^{-t/RC}$$

$$\ln Q = \ln Q_0 - \frac{t}{RC}$$

$$ln I = ln I_0 - \frac{t}{RC}$$

$$\ln V = \ln V_0 - \frac{t}{RC}$$

In a magnetic field

$$F = Bqv \sin \theta$$

$$F = BIl \sin \theta$$

Faraday's and Lenz's laws

$$\mathcal{E} = \frac{-\mathrm{d}(N\phi)}{\mathrm{d}t}$$

Nuclear and particle physics

In a magnetic field

$$r = \frac{p}{BQ}$$

Mass-energy

$$\Delta E = c^2 \Delta m$$



## Unit 5

**Thermodynamics** 

Heating  $\Delta E = mc\Delta\theta$ 

 $\Delta E = L\Delta m$ 

Ideal gas equation pV = NkT

Molecular kinetic theory  $\frac{1}{2}m < c^2 > = \frac{3}{2}kT$ 

Nuclear decay

Mass-energy  $\Delta E = c^2 \Delta m$ 

Radio-active decay  $A = -\lambda N$ 

$$\frac{\mathrm{d}N}{\mathrm{d}t} = -\lambda N$$

$$\lambda = \frac{\ln 2}{t_{1/2}}$$

$$N = N_0 e^{-\lambda t}$$

$$A = A_0 e^{-\lambda t}$$

Oscillations

Simple harmonic motion F = kx

 $a = -\omega^2 x$ 

 $x = A \cos \omega t$ 

 $v = -A\omega \sin \omega t$ 

 $a = A\omega^2 \cos \omega t$ 

$$T = \frac{1}{f} = \frac{2\pi}{\omega}$$

$$\omega = 2\pi f$$

Simple harmonic oscillator  $T = 2\pi \sqrt{\frac{m}{k}}$ 

$$T = 2\pi \sqrt{\frac{l}{g}}$$



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# Astrophysics and Cosmology

Gravitational field strength g = F/m

Gravitational force  $F = \frac{Gm_1m_2}{r^2}$ 

Gravitational field  $g = \frac{Gm}{r^2}$ 

Gravitational potential  $V_{grav} = \frac{-Gm}{r}$ 

Stephan-Boltzman law  $L = \sigma T^4 A$ 

Wein's law  $\lambda_{\text{max}} T = 2.898 \times 10^{-3} \text{ m K}$ 

Intensity of radiation  $I = \frac{L}{4\pi d^2}$ 

Redshift of electromagnetic  $z = \frac{\Delta \lambda}{\lambda} \approx \frac{\Delta f}{f} \approx \frac{v}{c}$  radiation

Cosmological expansion  $v = H_0 d$ 

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